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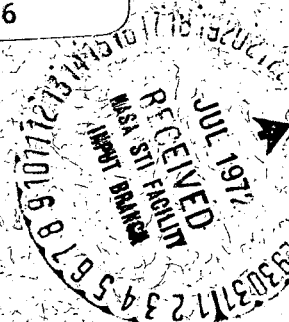
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ATS SIMULTANEOUS AND TURNAROUND RANGING EXPERIMENTS

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ABSTRACT

This report explains the data reduction and spacecraft position determination used in conjunction with two ATS experiments - Trilateration and Turnaround Ranging - and describes in detail a multilateration program that is used for part of the data reduction process. The process described is for the determination of the inertial position of the satellite, and for forming input for related programs. In the trilateration procedure, a geometric determination of satellite position is made from near simultaneous range measurements made by three different tracking stations. Turnaround ranging involves two stations; one, the master station, transmits the signal to the satellite and the satellite retransmits the signal to the slave station which turns the signal around to the satellite which in turn retransmits the signal to the master station. The results of the satellite position computations using the multilateration program are compared to results of other position determination programs used at Goddard. All programs give nearly the same results which indicates that because of its simplicity and computational speed the trilateration technique is useful in obtaining spacecraft positions for near synchronous satellites.

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ATS SIMULTANEOUS AND TURNAROUND RANGING EXPERIMENTS

I. INTRODUCTION

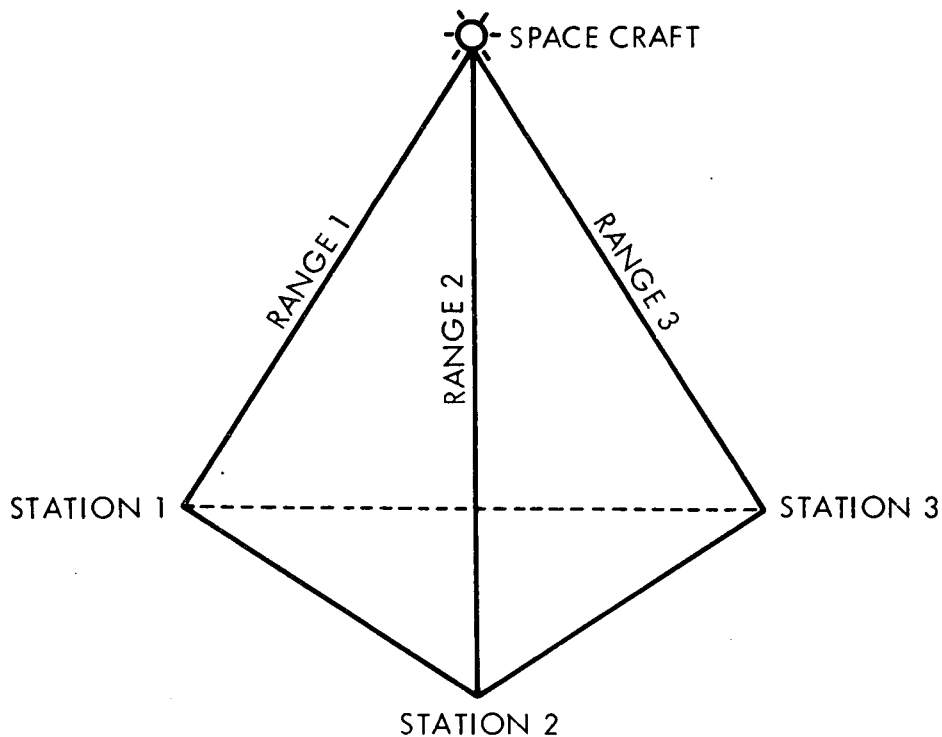
The purpose of this document is to explain the data reduction used in conjunction with two ATS spacecraft position determination experiments - Trilateration and Turnaround Ranging - and to describe in detail one of the programs used for these experiments. In the trilateration procedure, a geometric determination of satellite position is made from near simultaneous range measurements made by three different tracking stations (see Figure 1). Turnaround ranging involves two stations; one, the master station transmits the signal to the satellite and the satellite retransmits the signal to the slave station which turns the signal around to the satellite which in turn retransmits the signal to the master station.

The Trilateration experiment used near simultaneous measurements of range to obtain a position vector of the satellite. In reality, three stations took measurements during the same ten minute period of time and by fitting this data with three polynomials as a function of raw data time one is able to create groups of simultaneous measurements from the polynomials. See page 18 for details of the ranging pattern. These groups of simultaneous measurements were then fed into a Hughes trilateration program - Trilatron 1 (Reference 1), and out of each group of measurements one position vector of the satellite was obtained. The advantage of the trilateration technique is that it does not require a sophisticated force model that is required by conventional orbit determination programs. The trilateration method requires that the three tracking stations be located so as to give a well determined triangle, with the subsatellite point preferably being located in the triangle.

Turnaround Ranging uses a master station and some slave stations. The signal leaves the master station, travels to the satellite, returns to a slave station and is turned around to the satellite and then back to the master station, where it is recorded. By using this scheme, the slave station equipment need only be capable of turning the signal around to the satellite. Because the master station is tracking the satellite in its normal mode (two way delay time), the slave station range can be extracted from the four way delay time.

In addition to describing the Trilateration and Turnaround Ranging experiments and their results, this report describes a more general multilateration program to determine the position of a satellite in synchronous orbit. In this program the position of the satellite is computed by knowing the ranges from three or more observation points, and the times of the observations which should

be as near the same time as possible. To obtain a solution, the program begins by making an initial estimate of the inertial coordinates of the satellite. For the first observation time, this initial estimate is obtained by using the trilateration technique given in Reference 2. The initial estimate for each succeeding observation point is taken from the results of the previous one. From the inertial coordinates of the tracking station location and of the satellite, the program computes a range for each observation point and compares these to the observed ranges. By using a Taylor series expansion and a least squares process a converged set of the inertial coordinates of the satellite is then determined. If there are only three observation points, Cramer's rule may be used to solve the set of three simultaneous equations for the position of the satellite. In addition, the multilateration program referred to in this report may be used to obtain a solution if there are more than three observation points.



II. DATA TYPES

Two types of data were used. The two way delay data is the normal kind from the ATSR ranging system. It measures the time it takes for a signal to travel from the station to the satellite and return to the ground. This data type and its measurement is given in references 3 and 4. The other data type is four way delay time or Turnaround Ranging. This measurement signal leaves the master station, travels to the satellite, returns to the slave station, is turned around to the satellite and finally is returned to the master station. The master station must have tracked in the two way mode and in the same time period as well, in order to extract the slave station's range at this time from the turnaround range.

Trilateration Experiment

In order to be able to trilaterate in the sense of the Hughes trilateration program, it is necessary to obtain a series of overlapping range values to three different stations. The Hughes Trilateration program then creates a position and velocity vector from this information. For detailed information about this program see reference 1. One of the functions of the software package described in this document is to create the required format from the polynomials representing the raw data. See Reference 4 for raw data format information.

Turnaround Ranging

From the four way delay data a range to the satellite must be determined. The only station that needs to be fully equipped is the master station. The slave station need only to be able to turn the signal around. The output of this software package is a preprocessed series of ranges to be input to the trilateration program, and a DODS observation tape to be input to an orbit determination program.

III. ALGORITHM AND FORMULAS

This section will describe the formulas used and how they are derived to process this data. The details of the two way range data type are described in reference 3 and it is suggested that the reader read that document. The actual programmed equations will be marked with a starred number. It is hoped that the intermediate steps will make their derivation clear. The major problem in interpreting this data is that the time on the raw data message is not the proper observation time tag. Iterating so as to converge on the proper time tag and relating this time tag to the raw data time is the one complication of the procedure described in this algorithm.

A smoothing program creates a Chebyshev polynomial up to 12th degree of raw data and raw data time. See reference 3, Section 6.3. The spans of data have to be organized and read into this program in the groups that one wishes to have fitted with polynomials. There are sorting programs to do this. The polynomial coefficients and other details for creating the DODS Observation Tape are passed to the multilateration program where the ranges and proper time tags are then determined.

A time T_R , is chosen to be satellite time. The relationship between the raw data time and the ground received time is shown in the following figures. In order to find the proper measurement to correspond to the satellite time T_R , one needs the raw data time T_D that corresponds (the measurement polynomial is a function of raw data time). The connection between the times can be achieved by iterating to find the proper raw data time and therefore proper measurement to correspond to the satellite time, T_R . The convergence is very rapid and few iterations are necessary.

Let

T_D = raw data time

δ_2 = two way delay time raw measurement

D_M = total two way delay time of the master station

δ_4 = four way delay time raw measurement

D_4 = total four way delay time

D_S = total two way delay time of the slave station

N_{AM} = the ambiguity number for the master station

N_{AS} = the ambiguity number for the slave station

ΔA = the size of the ambiguity in time

T_R = the proper data time tag

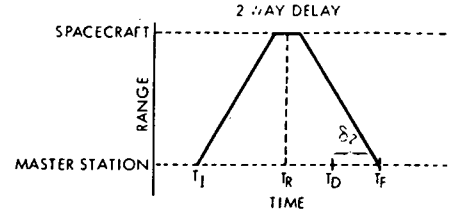
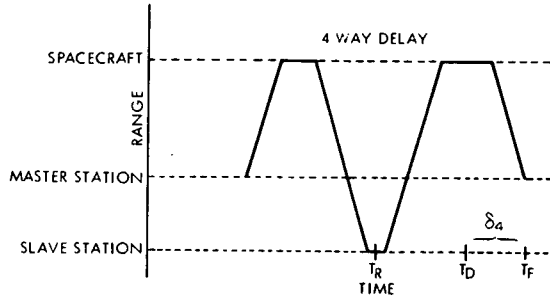
T_I = time the signal left the ground

T_F = time the signal returned to the ground

c = speed of light

S_M = master station delay

S_S = slave station delay



One first chooses sets of data which overlap in time. For both trilateration and turnaround ranging it is a requirement to have the ranges from all the stations involved at the same time.

1* T_R = first overlap time

Since the polynomials are in terms of T_D time an estimate of the T_D corresponding to T_R needs to be made. We will first consider the 2 way delay measurement.

Since (1)
$$T_R = \frac{T_I + T_F}{2}$$

(2)
$$T_I = T_D - N_{AM} \Delta A$$

(3)
$$T_F = T_D + \delta_2$$

(4)
$$T_R = T_D - \frac{N_{AM} \Delta A}{2} + \frac{\delta_2}{2}$$

(5)
$$T_D = T_R + \frac{N_{AM} \Delta A}{2} - \frac{\delta_2}{2}$$

and since δ_2 is unknown at this point the first approximation to T_D is

2*
$$T_D = T_R + \frac{N_{AM} \Delta A}{2} \quad \text{from equation (5)}$$

3* Then obtain δ_2 at T_D from the Chebyshev polynomial.

Compute from equation (4)

4*

$$T'_R = T_D - \frac{N_{AM} \Delta A}{2} + \frac{\delta_2}{2}$$

Compare

5* If $|T_R/T'_R - 1| > \epsilon$, T_R and T'_R have not converged. According to equation (5) T_D should be

$$T_D = T_R + \frac{N_{AM} \Delta A}{2} - \frac{\delta_2}{2}$$

and we approximated it by

2*

$$T_D = T_R + \frac{N_{AM} \Delta A}{2}$$

so let

6*

$$T_D = T_D - \frac{\delta_2}{2}$$

and try again from equation 3*.

5* If $|T_R/T'_R - 1| \leq \epsilon$, T_R and T'_R have converged, then compute

7* $D_M = \delta_2 + N_{AM} \Delta A - S_M$ the master station total delay.

If all the data is two way data one would repeat this process with each station.

At this point, one would have the proper measurement from each station at the same time. One could solve for a position vector at time T_R (trilaterate) if one has ranges from three stations.

If some of the data is four way (turnaround) data the following algorithm must be used.

The four way data contains the sum of ranges to two stations. It is necessary to extract the range to the slave station from this data. In order to do this, one must have tracked the satellite from the master station alone at approximately the same time. This is only necessary because the orbit determination programs currently do not accept the summed range as a data type. Due to this restriction, one relies on the smoothing polynomial to fill in the gaps in the data. It is assumed at this point that the master station total delay is known at time T_R , as obtained in the previous algorithm. It is now necessary to find the four way total delay, D_4 and then to compute the total two way delay for the slave station ($D_S = D_4 - D_M$).

Choose the same T_R as before.

Since (6)
$$T_R = \frac{T_I + T_F}{2}$$

(7)
$$T_I = T_D - N_{AM} \Delta A - N_{AS} \Delta A$$

(8)
$$T_F = T_D + \delta_4$$

(9)
$$T_R = T_D - \left(\frac{N_{AM} \Delta A + N_{AS} \Delta A}{2} \right) + \frac{\delta_4}{2}$$

(10)
$$T_D = T_R + \left(\frac{N_{AM} \Delta A + N_{AS} \Delta A}{2} \right) - \frac{\delta_4}{2}$$

and since δ_4 is unknown at this point the first approximation to T_D is

8*
$$T_D = T_R + \left(\frac{N_{AM} \Delta A + N_{AS} \Delta A}{2} \right).$$

9* Then obtain δ_4 at T_D from the Chebyshev polynomial.

Compare δ_4 and δ_2 .

If $\delta_4 \leq \delta_2$ then an extra ambiguity had to occur because of the combined effect of two stations. This means equation (10) needs an extra $\Delta A/2$ added to it and

$$(11) \quad T_D = T_R + \left(\frac{N_{AM} \Delta A + N_{AS} \Delta A + \Delta A}{2} \right) - \frac{\delta_4}{2}$$

for this case.

Therefore

$$10^* \quad T_D = T_D + \frac{\Delta A}{2}$$

and since the initial T_D had to be far off another $10^* \delta_4$ is obtained from the Chebyshev polynomial at the new T_D .

D_4 is then computed.

$$11^* \quad D_4 = \delta_4 + N_{AM} \Delta A + N_{AS} \Delta A + \Delta A - S_M - S_S$$

$$12^* \quad T'_R = T_D - \left(\frac{N_{AM} \Delta A + N_{AS} \Delta A + \Delta A}{2} \right) + \frac{\delta_4}{2} \text{ from equation } (11)$$

If the new $\delta_4 > \delta_2$ recompute T_R and D_4 as if that were the original situation as below. When $\delta_4 > \delta_2$ compute

$$13^* \quad T'_R = T_D - \left(\frac{N_{AM} \Delta A + N_{AS} \Delta A}{2} \right) + \frac{\delta_4}{2} \text{ from equation } (9)$$

$$\text{and } 14^* \quad D_4 = \delta_4 + (N_{AM} \Delta A + N_{AS} \Delta A) - S_M - S_S$$

If $15^* \quad |T_R/T'_R - 1| \leq \epsilon$, T_R and T'_R have converged then

$$16^* \quad D_S = D_4 - D_M$$

If $15^* \quad |T_R/T'_R - 1| > \epsilon$, T_R and T'_R have not converged then

$$\text{let } 16^* \quad T_D = T_D - \frac{\delta_4}{2}$$

and try again from equation 3^* .

After two way delays are determined for each station the range R for each station is computed

$$17^* \quad R = \frac{c}{2} D, \text{ where } D = D_M \text{ or } D_S$$

The output is properly formatted and the data is time tagged, thus completing this task.

IV. COMPUTATION OF THE INERTIAL COORDINATES OF THE TRACKING STATIONS

From the results of the previous section, we now have a set of ranges (three or more) at the same time. The function of the multilateration program is to compute the sub-satellite point at this time.

The multilateration method takes the set of three or more simultaneous ranges, uses a least squares solution to compute the inertial satellite position, and then computes the sub-satellite point in the earth fixed system. These items are computed according to the formulas presented in sections IV, V, and VI.

However, first one needs to know the coordinates of the ranging stations. The method for computing the inertial coordinates of the tracking stations is the one presented on pages 16-18 of Reference 5.

The following data are needed to compute the coordinates of the tracking stations:

- (λ_0) = the hour angle of the first point of Aries
- (λ_E) = the geodetic longitude of the terrestrial tracking station in radians, as measured eastward from Greenwich (a negative sign must be prefixed if measured westward from Greenwich)
- (θ_D) = the geodetic latitude of the station in radians, measured as positive north of the Equator, and as negative south of the Equator
- (H') = the altitude of the station in feet, measured positive above sea level and negative below sea level
- (ω) = the angular velocity of rotation of the earth in radians per hour
- (ΔT) = the difference in hours between the observation time and midnight preceding the observation time

- (f) = the flattening coefficient of the earth
- (θ_G) = the geocentric latitude of the tracking station in radians
- ($\hat{\rho}$) = the geocentric distance of the tracking station, in units of earth radii
- (e) = the eccentricity of the earth

The geodetic longitude (λ_E), geodetic latitude (θ_D), and height (H) of the tracking station are known. From these the inertial geocentric coordinates of the tracking station in spherical coordinates ($\hat{\rho}$, θ_G , δ) can be computed. These spherical coordinates can then be converted to a Cartesian system of coordinates (x_T , y_T , z_T).

In the meridian section of the earth through an observer, the position of the latter relative to the earth's center can be expressed in rectangular coordinates as:

$$\hat{\rho} \sin \theta_G = S \sin \theta_D, \quad \hat{\rho} \cos \theta_G = C \cos \theta_D.$$

These serve to define the auxiliary functions S and C.

$$C = [\cos^2 \theta_D + (1 - f)^2 \sin^2 \theta_D]^{-1/2}$$

$$S = (1 - f)^2 C$$

$$H = (4.77865 \times 10^{-8}) H' \text{ (converts feet to earth radii)}$$

$$\theta_G = \arctan \left[\left(\frac{S + H}{C + H} \right) \right] \tan \theta_D$$

$$\hat{\rho} = [(S + H)^2 \sin^2 \theta_D + (C + H)^2 \cos^2 \theta_D]^{1/2}$$

$$\delta = \lambda_0 + \omega(\Delta T) + \lambda_E$$

- (δ) = the angle in radians between the vernal equinox and the observation meridian plane

$$x_T = \hat{\rho} \cos \theta_G \cos \delta$$

$$y_T = \hat{\rho} \cos \theta_G \sin \delta$$

$$z_T = \hat{\rho} \sin \theta_G$$

(x_T, y_T, z_T) = the geocentric coordinates of the tracking station in units of earth radii

We now have the station locations in quantities and units that we can use in the program.

V. COMPUTATION OF THE INERTIAL COORDINATES OF THE SATELLITE

This multilateration method computes the inertial coordinates of the satellite by a least squares iteration procedure, given the ranges from the satellite to three or more tracking stations, the inertial coordinates of the tracking stations, and an initial estimate of the inertial coordinates of the satellite. This section describes the mathematical method for computing the inertial coordinates of the satellite. We refer to this method as the multilateration method.

$(R_o)_i$ = observed range in earth radii

(R_c) = computed range in earth radii

(x, y, z) = the present estimate of the inertial coordinates of the satellite, obtained initially by using the trilateration method given in Reference 2.

(x', y', z') = the new estimate of (x, y, z)

$$R_c^2 = (x - x_T)^2 + (y - y_T)^2 + (z - z_T)^2$$

A function of three variables may be expanded in a series by Taylor's formula in the form:

$$f(x + \Delta x, y + \Delta y, z + \Delta z) - f(x, y, z)$$

$$= \frac{\partial f}{\partial x} \Delta x + \frac{\partial f}{\partial y} \Delta y + \frac{\partial f}{\partial z} \Delta z$$

+ higher order terms

or $\Delta f = \frac{\partial f}{\partial x} \Delta x + \frac{\partial f}{\partial y} \Delta y + \frac{\partial f}{\partial z} \Delta z$, ignoring the higher order terms

since $R_C^2 = f(x, y, z)$

we can write
$$\Delta R^2 = \frac{\partial R_C^2}{\partial x} \Delta x + \frac{\partial R_C^2}{\partial y} \Delta y + \frac{\partial R_C^2}{\partial z} \Delta z$$

where $\Delta R^2 = R_O^2 - R_C^2$

$$\frac{\partial R_C^2}{\partial x} = 2x - 2x_T$$

$$\frac{\partial R_C^2}{\partial y} = 2y - 2y_T$$

$$\frac{\partial R_C^2}{\partial z} = 2z - 2z_T$$

Therefore
$$\Delta R^2 = (2x - 2x_T) \Delta x + (2y - 2y_T) \Delta y + (2z - 2z_T) \Delta z$$

or
$$\frac{1}{2} \Delta R^2 = (x - x_T) \Delta x + (y - y_T) \Delta y$$

$$+ (z - z_T) \Delta z \tag{5.1}$$

A least squares routine is then used to compute $\Delta x, \Delta y, \Delta z$. The least squares method used is the one described on pages 61-66 of Reference 5.

Then

$$x' = x + \Delta x$$

$$y' = y + \Delta y$$

$$z' = z + \Delta z$$

The program then returns to Equation 5.1 with x, y, z being replaced by $x', y',$ and z' . ΔR^2 is re-evaluated. This iterative procedure is continued until the standard deviation of fit of $1/2 \Delta R^2$ reaches the desired tolerance, and we thus have the inertial position of the spacecraft.

VI. COMPUTATION OF THE LONGITUDE AND LATITUDE OF THE SUB-SATELLITE POINT

The method for transforming the inertial position coordinates (x, y, z) of the satellite, as computed above, to the East longitude (λ_s), the geodetic latitude (θ_D), and the elevation of the satellite (H_s) above and normal to the adopted ellipsoid, and the geocentric radius magnitude (ρ) is given in this section.

The symbols used in this section are the same as those used in Section IV, except they refer to the sub-satellite point instead of the tracking station location. The method used in this section may be found in Reference 1.

The converged values of the geocentric coordinates of the satellite (x, y, z) have been found using Section V. Proceeding from these knowns, we compute the unknown longitude, latitude, and elevation as follows:

$$\rho^2 = x^2 + y^2 + z^2$$

$$e^2 = 2f - f^2$$

$$\cos \theta_G = \left[\frac{(x^2 + y^2)}{\rho^2} \right]^{1/2}$$

$$\sin \theta_G = \frac{z}{\rho}$$

$$\tan \theta_G = \frac{\sin \theta_G}{\cos \theta_G}$$

$$\tan \theta_G = (1 - f)^2 \tan \theta_D$$

$$\theta_D = \arctan \left[\frac{\tan \theta_G}{(1 - f)^2} \right]$$

$$\cos \delta = \frac{x}{\rho \cos \theta_G}$$

$$\sin \delta = \frac{y}{\rho \cos \theta_G}$$

$$\lambda_s = \delta - \lambda_0 - \omega(\Delta T)$$

if λ_s is greater than 180° , the program will take λ_s as

$$\lambda_s = \lambda_s - 360^\circ$$

$$\Delta \theta'_G = 0$$

$$\theta'_G = \theta_G - \Delta \theta'_G \quad (6.1)$$

$$r_c = \left[\frac{1 - e^2}{1 - e^2 \cos^2 \theta'_G} \right]^{1/2}$$

$$\theta_D = \tan^{-1} \left[\frac{\tan \theta'_G}{(1-f)^2} \right]$$

$$H = [\rho^2 - r_c^2 \sin^2 (\theta_G - \theta'_G)]^{1/2} - r_c \cos (\theta_G - \theta'_G)$$

$$\Delta\theta'_G = \sin^{-1} \left[\frac{H}{\rho} \sin (\theta_G - \theta'_G) \right]$$

If $\Delta\theta'_G$ has not stopped varying return to Equation (6.1).

$$H_S = a_e H$$

(λ_s) = the geodetic longitude of the sub-satellite point (+ indicates east of Greenwich and - west of Greenwich)

(θ_D) = the geodetic latitude of the sub-satellite point (+ indicates north of the Equator and - south of the Equator)

(a_e) = radius of the earth in kilometers

(H_S) = the height of the satellite above and normal to the adopted ellipsoid.

Thus we have computed the longitude and latitude of the sub-satellite point, and the height of the satellite.

VII. RESULTS

Simulated Data

The multilateration program described herein has been tested using simulated data for the ATS-1 satellite. The epoch used was the 4th of April 1969 at 0 hours, 0 minutes, and 0 seconds.

Some of the orbital parameters were:

semi-major axis = 42,166.5 km.

eccentricity = .000229

inclination = 1.4°

period = 23.94 hr.

The test cases listed below were run using the three observations given below.

<u>Tracking Station</u>	<u>Observation Time</u>						<u>Range</u>
	<u>yr.</u>	<u>mo.</u>	<u>day</u>	<u>hr.</u>	<u>min.</u>	<u>sec.</u>	
Mojave, Calif.	69	4	4	0	37	0	38,133.6887 km.
Rosman, N. C.	69	4	4	0	37	0	40,633.4199
Toowoomba, Australia	69	4	4	0	37	0	39,518.6607

The true values of x, y, z as given by the program which simulated the data were:

$$x = 26,297.9477 \text{ km.}$$

$$y = 32,944.5272$$

$$z = -573.1739$$

The initial values used for the three test cases were:

	<u>x</u>	<u>y</u>	<u>z</u>
Case 1	+28,701.7425 km.	+28,701.7425 km.	-4464.7155 km.
Case 2	+19,134.4950	+22,323.5775	-6314.3833
Case 3	+25,512.6600	+31,890.8250	-574.0348

All the cases converged to the same values:

$$x = 26,298.1470 \text{ km.}$$

$$y = 32,944.3690$$

$$z = -573.1558$$

If we compare these values to the given values, we get:

$$\Delta x = -200 \text{ m.}$$

$$\Delta y = +158$$

$$\Delta z = -18$$

Where the difference is taken as the given value minus the computed value.

The sub-satellite point is then found to be:

$$\begin{aligned}\text{Longitude} &= -141.66^\circ \\ \text{Latitude} &= -0.784^\circ \\ \text{Height of satellite} &= 35,779.31 \text{ km.}\end{aligned}$$

The program gave similar results when Cramer's method was used instead of a least squares procedure. For example, using Cramer's method for Case 1 gave the following results:

$$\begin{aligned}x &= 26,298.1546 \text{ km.} \\ y &= 32,944.3582 \\ z &= -573.1764 \\ \Delta x &= -207 \text{ m.} \\ \Delta y &= +169 \\ \Delta z &= +2 \\ \text{Longitude} &= -141.660^\circ \\ \text{Latitude} &= -0.784^\circ \\ \text{Height of satellite} &= 35,779.92 \text{ km.}\end{aligned}$$

The largest error in the test cases was 207 meters. These test cases do not agree exactly because the program generating the simulated data and the multilateration program did not use the same sidereal time. These test cases were run before the trilateration method for computing the initial estimates of x , y , z had been incorporated into the multilateration program. Guided by these results, the program was then run for three cases using real satellite data.

Real Data

Case 4

The observations for this case were taken on October 1, 1969 by three ATS ground stations: Rosman, North Carolina; Mojave, California; and Toowoomba, Australia. Mojave was the prime or two-way delay station; whereas, Toowoomba and Rosman were four-way delay stations. Two-way delay means that the signal is sent from the station to the satellite and is then returned to the station. Four-way delay means that the signal is sent from station A to the spacecraft, from the spacecraft to station B, from station B to the satellite, and from the satellite to station A. Station A would be the master station in this case.

The three stations tracked the spacecraft during a ten minute span of time as illustrated in Figure 2. The data rate was one observation per second and each station ranged for about one minute each time it tracked.

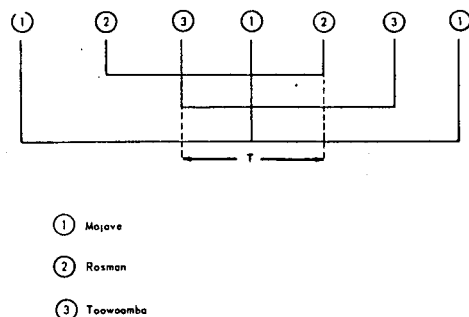


Figure 2.

Some of the orbital parameters used are:

semi-major axis = 42,168.1 km.
eccentricity = .00229
period = 23.94 hr.

The initial values used for the inertial coordinates are:

x = +5740.3485 km.
y = -41,458.0725
z = -3507.9907

The results are then compared to the results obtained by the Refined World Map Program (Reference 6).

<u>Program</u>	<u>Observation Time</u>			<u>Satellite Position</u>		<u>HT</u>
	<u>hr.</u>	<u>min.</u>	<u>sec.</u>	<u>Long. (+E)</u>	<u>Lat. (+N)</u>	
ATS	3	47	56.00	-148.710°	-.484°	35782.0 km.
WMAP	3	48	0.00	-148.650	-.489	35787.0
ATS	3	53	56.01	-148.734	-.536	35780.7
WMAP	3	54	0.00	-148.655	-.537	35788.7
WMAP	4	41	0.00	-148.660	-.896	35789.0
ATS	4	41	35.00	-148.779	-.822	35787.0
WMAP	4	42	0.00	-148.660	-.903	35790.0

Here, ATS refers to the multilateration program developed in this report.

The longitude agrees to within $.1^\circ$, the latitude within $.01^\circ$ and the height within 10 kilometers. Some of the error is probably due to the delays in the ground equipment at the master station not being measured accurate enough.

Case 5

The observations for this case were taken on December 23, 1969 by the ATS ground stations: Rosman, Mojave, and Toowoomba. This time the stations made simultaneous observations. All the stations were therefore two-way delay stations. They give an initial estimate for x, y, z of

$$\begin{aligned}x &= +21,934.0285 \text{ km.} \\y &= -35,990.6776 \\z &= -1103.2987\end{aligned}$$

Below is the comparison of the results of the multilateration program (called ATS in the charts) with the ORB-1 and WMAP programs (Reference 6).

Observation Time

<u>Program</u>	<u>yr.</u>	<u>mo.</u>	<u>day</u>	<u>hr.</u>	<u>min.</u>	<u>sec.</u>	<u>x</u>	<u>y</u>	<u>z</u>
ATS	69	12	23	0	22	0.02	25,677.7275 km.	-33,421.9067 km.	-1124.1973 km.
ORB-1	69	12	23	0	22	0.00	25,683.7708	-33416.8086	-1125.5546
ATS	69	12	23	0	23	1.02	25,826,0626	-33,307.2853	-1128.8897
ORB-1	69	12	23	0	23	0.00	25,829.6496	-33,304.0089	-1130.0527

Observation Time

Satellite Position

<u>Program</u>	<u>yr.</u>	<u>mo.</u>	<u>day</u>	<u>hr.</u>	<u>min.</u>	<u>sec.</u>	<u>Long. (+E)</u>	<u>Lat. (+N)</u>	<u>Ht</u>
ATS	69	12	23	0	22	0.02	-149.340°	-1.538°	35783.8 km.
WMAP	69	12	23	0	22	0.00	-149.327	-1.531	35784.0
ATS	69	12	23	0	23	1.02	-149.340	-1.545	35783.8
WMAP	69	12	23	0	23	0.00	-149.327	-1.537	35784.0

The longitude agrees to within $.1^\circ$ the latitude within $.01^\circ$, and the height within 1 kilometer.

Case 6

The observations for this case were taken on February 20 and 21, 1970. The observations were taken simultaneously by the three tracking stations of Mojave, Kashima, and Toowomba.

Below is a comparison of the multilateration program (called ATS in the charts), with the TV-1, ORB-1, and WMAP programs.

Program	<u>Observation Time</u>						<u>Satellite Position</u>		
	yr.	mo.	day	hr.	min.	sec.	<u>x</u>	<u>y</u>	<u>z</u>
ATS	70	2	20	4	40	30.26	+14843.66km.	+39476.84km.	-349.72km.
TV-1	70	2	20	4	40	30.26	+14844.79	+39476.42	-349.72
ORB-1	70	2	20	4	40	30.00	+14844.87	+39476.33	-350.27
ATS	70	2	20	10	40	29.26	-39487.29	+14695.89	+1642.31
TV-1	70	2	20	10	40	29.26	-39486.87	+14702.83	+1642.26
ORB-1	70	2	20	10	40	30.00	-39487.87	+14699.94	+1641.75
ATS	70	2	20	17	10	30.25	-9206.24	-41136.44	+120.00
TV-1	70	2	20	17	10	30.25	-9207.42	-41136.18	+120.00
ORB-1	70	2	20	17	10	30.00	-9207.20	-41136.14	+119.27
ATS	70	2	20	21	10	28.28	+31088.54	-28439.97	-1391.48
TV-1	70	2	20	21	10	29.29	+31087.26	-28438.56	-1391.55
ORB-1	70	2	20	21	10	30.00	+31096.58	-28436.10	-1392.39

Program	<u>Observation Time</u>						<u>Satellite Position</u>		
	yr.	mo.	day	hr.	min.	sec.	<u>Long. (+E)</u>	<u>Lat. (+N)</u>	<u>Ht.</u>
ATS	70	2	20	4	40	30.26	-150.44°	- .48°	35798.6 km.
TV-1	70	2	20	4	40	30.26	-150.44	- .48	35798.6
WMAP	70	2	20	4	40	30.00	-150.43	- .47	35798.4
ATS	70	2	20	10	40	29.26	-150.49	+2.23	35789.2
TV-1	70	2	20	10	40	29.26	-150.49	+2.23	35789.2
WMAP	70	2	20	10	40	30.00	-150.49	+2.23	35789.0
ATS	70	2	20	17	10	30.25	-150.46	+0.16	35776.0
TV-1	70	2	20	17	10	30.25	-150.46	+0.16	35776.0
WMAP	70	2	20	17	10	30.00	-150.45	+0.16	35775.8
ATS	70	2	20	21	10	28.28	-150.45	-1.89	35778.5
TV-1	70	2	20	21	10	29.29	-150.45	-1.89	35778.5
WMAP	70	2	20	21	10	30.00	-150.45	-1.89	35778.4

The longitude agrees to within 0.1° , the latitude within $.01^\circ$, and the height within 1 kilometer. The TV-1 and the multilateration programs did not give the same results, because the TV-1 programs used mean sidereal time and the multilateration program used apparent sidereal time.

From these results it can be seen that there is excellent agreement between the various position determination methods. The trilateration method has advantages over conventional orbit determination techniques in that it is a geometric solution and does not require a sophisticated force model, nor does it require a number of iterations to obtain an orbit as do conventional orbit determination techniques. Each position vector takes less than five seconds CPU time on the 360/91 when using the trilateration program. The time required by conventional orbit determination programs depends on the length of the arc needed to determine the orbit, the number of iterations necessary to converge, and the force model used. Some disadvantages of the described trilateration technique are; an orbit is not determined and therefore orbit prediction can not be made, and there must be mutual visibility by three stations capable of tracking almost simultaneously the particular satellite. For additional information on the future development of the trilateration technique the reader should refer to references 8 and 9.

VIII. CONCLUSION

From the results of the cases given in the previous section, it can be seen that the multilateration programs give good results. The Multilateration, TV-1, ORB-1 and WMAP programs compared very favorably. The longitude always agreed to within $.1$ degrees, and the latitude always agreed to within $.01$ degrees. The height agreed to within 1 kilometer, except on the turnaround ranging case. A possible explanation for this discrepancy is that the measurements of the time delay in the ground equipment at the slave station is inaccurate. The results further indicate that spacecraft positions for synchronous satellites can be obtained using less computer time and less computer memory if one uses the multilateration instead of using conventional orbit determination techniques.

IX. ACKNOWLEDGMENT

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APPENDIX
Program Information

A. Operating Instructions for Multilateration Program

Below is an explanation of the data cards, with some typical values in parenthesis

1. Card #1 - 5 variables in the format (5E15.8)
 - a. f = the flattening coefficient of the earth (.33528919)
 - b. TOL = the tolerance that is used to determine when the least squares program has converged to a solution ($.1 \times 10^{-8}$)
 - c. DELA = the size of the ambiguity in time in seconds (.125)
 - d. REARTH = the radius of the earth in kilometers (6378.165)
 - e. C = the speed of light in km./sec. (2.997925×10^5)
2. Card #2 - 7 variables in the format (2(I5, 5X), F5.1, I5, 5X, E15.8, 2I5)
 - a. NOOBS = the number of tracking stations being used in this run. The program with slight modification will handle up to 10 stations.
(3)
 - b. NOPRST = the station number of the prime or two way delay station.
If all stations are two way delay stations this number would be set to -1. (47)
 - c. DELTAT = the time increment in minutes at which the data points will be printed out in each time interval. (1)
 - d. NOPTS = the number of data points that are to be computed in each time interval. This option would override the value of DELTAT. This option is usually not used and the value is set to -1. (-1)

- e. EPSLON = the tolerance which is used to determine when the program has converged on the proper time (1.0×10^{-18}).
 - f. IMON = the month in which the observations begin (10).
 - g. IDAY = the day of the month on which the observations begin (1)
3. Card #3 to card # (2 + NOOBS) - 3 variables in the format (2I5, 5X, E15.8))
- a. NUMST = the number of the tracking station (47).
 - b. NA = the number of ambiguities (2)
 - c. STNDEL = the station delay time in micro-seconds (1.75).
4. Card #(2 + NOOBS) to Card # (32 + NOOBS) - 1 variable in the format (7X, E15.8). 30 cards total
- a. ALAMDM = the hour angle of the first point of Aries at midnight Greenwich mean time in radians. The first card in this set contains the hour angle for the IMON and IDAY on card #2. Each subsequent card has the hour angle for the next day. As the program is presently set up 30 cards are required, however if there is data only on one day only the first card need be filled in the other 20 may be blank. If NOOBS on card #2 is 2 or less all 30 cards may be blank. (.16668687, value for Oct. 1, 1969, obtained from Reference 8.)

Last 4 cards give the coordinates of the four ATS tracking stations in the format (13X, E15.8, 1X, E15.8, 1X, F8.1).

- a. ALAMDE = the geodetic longitude of the tracking station in radians as measured positive eastward from Greenwich (+4.2430894)
- b. THETAD = the geodetic latitude of the tracking station in radians as measured positive north of the equator (+.61665814).
- c. ALT = the altitude of the tracking station in feet measured as positive above sea level (+3072.8).

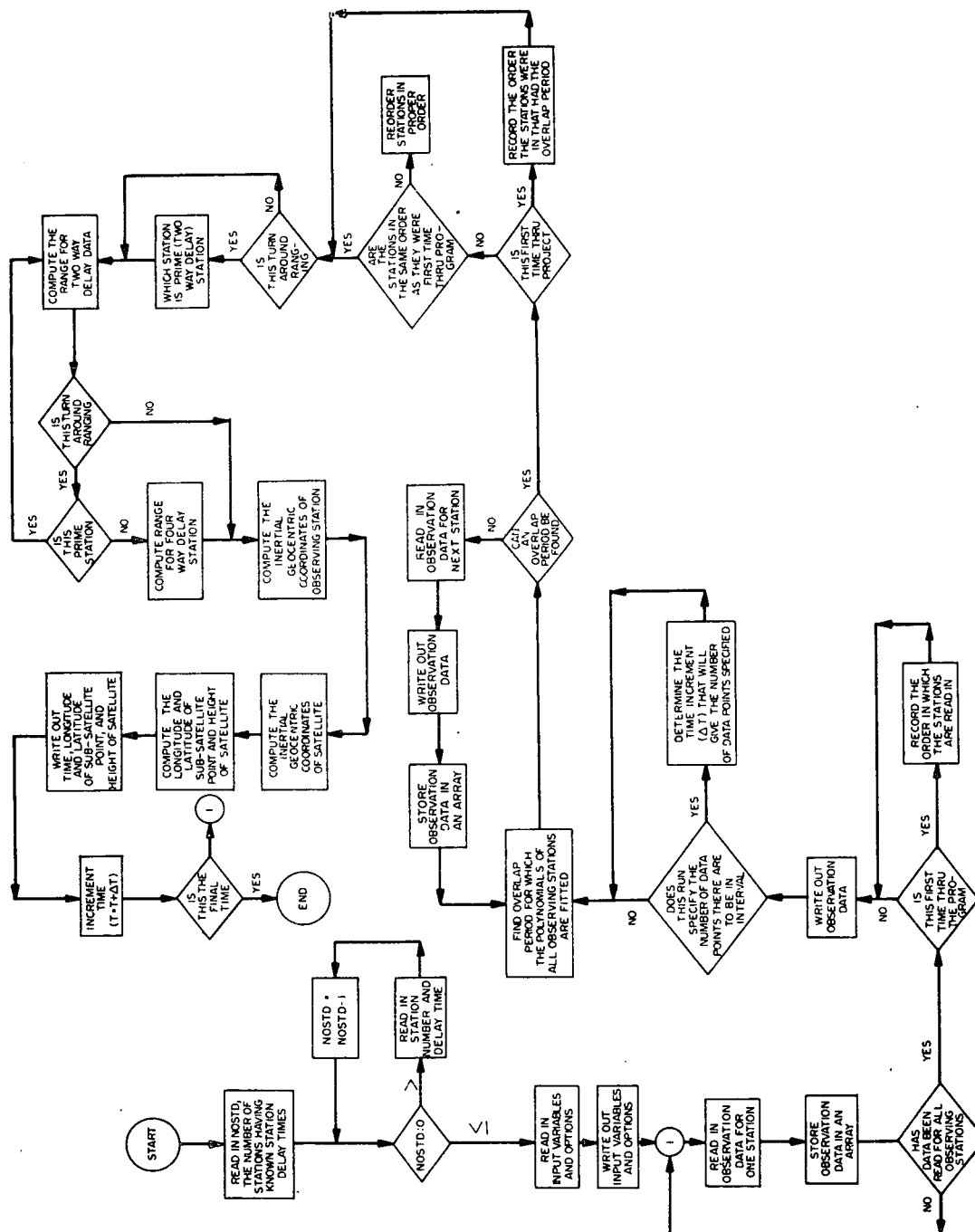
Note: The name of the station is given in columns 3-8 (Mojave in the example given in a thru c above), but is not used in the program.

B. Program Flow Chart

The program is presently set up for the ATS satellites and must have the four tracking stations in the order given below.

1. Kashima, Japan
2. Mojave, California
3. Rosman, North Carolina
4. Toowoomba, Australia

If any other stations but these four stations are used the multilateration program has to be slightly modified, because the station numbers are hard coded in the program.



C. Multilateration Program

```

IMPLICIT REAL*8(A-H,O-Z,$)
REAL MAXR,MAXRD,MAXXY,MAX,MIN
INTEGER CC1,CC2,CC3,CC4,C1SAV,C2SAV,C3SAV,C4SAV,U,V
REAL*8 OTIME,STID,OBSV,TRANOM,FICRP,OBCV
INTEGER*4 ISATN,IOBN
REAL*4 OBWT,COC,POC,SOF1,FNRRES,TMER,TCRV,ROBWT,RROBWT,AOBWT
INTEGER*2 IESN,IDRPN,IOT,IEF,ITTI,ITCI,IOCI
DIMENSION A(10,13), AC(13), ALAMCM(30), CLMQI(10,3), COEFF(13),
1 DELAY(10), DELR(10,1), CLTAQI(3,4), END(10), NDAY(12), NDEG(10),
2 NCMST(10), R(90), RCBS(10), STACCN(4,3), START(10), TYPE(10),
3 WT(10), XT(10), YT(10), ZT(10), RTVI(200,3), TVTIME(200,3),
4 I11(2), STATNM(10), STDFIT(10), WTOB(10), DELOUT(200,10),
5 NSTAT(10), STOR(13), STADEL(100), DELDOD(200,10), NAMB(100)
DIMENSION NAMBPR(10)
EQUIVALENCE (FIDRP,I11)
NAMelist/NAM1/N, NUMST, Y, NAMB, DELA, STNDEL, DEL2W, D2WAY, DELAY
1 , Y2, Y4
NAMelist/NAM2/N, NUMST, Y, NAMB, DELA, STNDEL, DEL2W, D2WAY, DELAY
1 , Y2, Y4
NAMelist/NAM3/N, NUMST, Y, NAMB, DELA, STNDEL, DEL2W, D2WAY, DELAY
1 , Y2, Y4
C CORRECTION FOR LIGHT TIME AND AMBIGUITY REQUESTED
INTEGER*2 IAMBRS
DIMENSION IAMBRS(3)
DATA IAMBRS/16,22,31/
C RANGE RR ANGLES CORRECTIONS TO BE MADE
DATA TRANOM,FIDRP,OBSV,CBWT,COC,POC,FNRRES,TMER,TCRV,IESN,IDRPN,
1 IOT,IEF,ITTI,ITCI,IOCI/5*0.,7*0/
196 IOT = 1
AE = 6378.165
RADCEG = 57.2957795
NDAY(1) = 31
NDAY(2) = 28
NDAY(3) = 31
NDAY(4) = 30
NDAY(5) = 31
NDAY(6) = 30
NDAY(7) = 31
NDAY(8) = 31
NDAY(9) = 30
NDAY(10) = 31
NDAY(11) = 30
NDAY(12) = 31
C
C IF NCPTS = 1 OR > 1, THE PROGRAM WILL COMPUTE A RANGE FOR THE
C NOPTS SPECIFIED. NO COMPUTATION OF INERTIAL COORDINATES OR
C OF SUB-SATELLITE POINT WILL BE MADE IN THIS CASE.
C NCPRST = -1 FOR THIS CASE. SET NOPTS = 0 FOR REGULAR CASE.
C IN THIS CASE THE PROGRAM WILL COMPUTE THE TIME COMMON TO ALL
C STATIONS AND COMPUTE RANGE, AND SUB-SATELLITE POINT FOR THIS TIME
C AND CONTINUE TO COMPUTE AT SPECIFIED INTERVALS OF TIME.
C
READ(5,199) F, TOL, DELA, REARTH, C, NOOBS, NCPRST, DELTAT, NOPTS,
1 EPSLON, IMON, IDAY
199 FORMAT(5E15.8/2(15,5X), F5.1, 15, 5X, E15.8, 215)
WRITE(6,5000) F, TOL, DELA, NOOBS, NCPRST, DELMIN, REARTH, C
WRITE(6,5019)
5C19 FORMAT(1H0, 37X, 17#TRACKING STATIONS)
DO 195 I = 1, NOOBS
READ(5,194) NUMST, NA, STNDEL
194 FORMAT(215, 5X, E15.8)
STNDEL = STNDEL * 1.0D-06
WRITE(6,5020) NUMST, NA, STNDEL

```

```

5C20 FORMAT(1H0,
117H NO. OF STATION =, I5, 5X, 14H NO. OF AMB. = I5, 5X,
216H STATION DELAY =, E20.8, 2X, 4PSEC.//)
STADEL(NUMST) = STNCEL
195 NAMB(NUMST) = NA
DELMIN = DELTAT

C
C IF NOPRST = -1, ALL STATIONS ARE TWO WAY DELAY STATIONS
C
C READ(5,200) (ALAMDM(I), I = 1, 30)
200 FORMAT(7X, E15.8)

C
C READ IN TRACKING STATION LOCATIONS
C
201 FORMAT(13X, E15.8, 1X, E15.8, 1X, F8.1)
READ(5,201) ((STACCN(I,J), J = 1,3), I = 1,4)
IOBN = 0
TEMP2 = DELA / 86400.0
DELTAT = DELTAT / 1440.0
N = 1
NOTIM = 0
NOITER = 1
III = 0
NOREAD = 0
JJ = 1

IF(NOPTS) 3, 3, 4
3 NK = NOOBS
GO TO 5
4 NK = 1
5 DO 10 I = 1, NK
NOREAD = NOREAD + 1
INDIC = -1
INDCT = -1
REAC(19,END=8998) STTIM, FINTIM, NUMST, DATYP, NDGREE, COEFF,
11SAT, ROBWT, SDFT, STNAME, FIDRP, IOCI

C
C THIS MEANS NO GATING IS DONE BY DODS, NOTICE THIS IS GOOD FOR
C RANGE ONLY, DO NOT USE IF RANGE RATE IS ALSO USED
C

I11(2) = I11(2) + 2
NOMST(I) = NUMST
IF(NOREAD - NOOBS) 6, 6, 7
6 III = III + 1
NSTAT(III) = NOMST(I)
7 START(I) = STTIM
END(I) = FINTIM
TYPE(I) = CATYP
NDEG(I) = NDGREE
STATNM(I) = STNAME
STDFIT(I) = SDFT
WTOB(I) = ROBWT
KK = NDEG(I) + 1
DO 11 J = 1, KK
11 A(I,J) = COEFF(J)
12 IF(J - 13) 13, 10, 10
13 J = J + 1
A(I,J) = 0.
GO TO 12
10 WRITE(6,5001) I, START(I), END(I), NOMST(I), TYPE(I), NDEG(I),
1(A(I,J), J = 1, 13)

```



```

5001 FORMAT(1H0, 3H1 = 13//
110H START = E20.8//10H END = E20.8//10H NOMST = 15//
210H TYPE = E20.8//10H NDEG = 15 //10H A(1) = E20.8//
310H A(2) = E20.8//10H A(3) = E20.8//10H A(4) = E20.8//
410H A(5) = E20.8//10H A(6) = E20.8//10H A(7) = E20.8//
510H A(8) = E20.8//10H A(9) = E20.8//10H A(10) = E20.8//
610H A(11) = E20.8//10H A(12) = E20.8//10H A(13) = E20.8//
GO TO 215
139 INDOCT = +1
IF(NOPTS) 415, 415, 140
140 M = 1
ENDTIM = FINTIM
IF(NCPTS - 2) 141, 142, 143.
141 T = STTIM + ((ENDTIM - STTIM) / 2.)
DELTAT = ENDTIM
GO TO 25
142 DELTAT = (FINTIM - STTIM) / 3.
DELTAT = DELTAT - TEMP3
T = STTIM + DELTAT
ENDTIM = FINTIM - (DELTAT / 2.)
GO TO 25
143 DELTAT = (ENDTIM - STTIM) / (NOPTS - 1.)
DELTAT = DELTAT - TEMP3
T = STTIM
GO TO 25
415 BEGTIM = START(1)
ENDTIM = END(1)
DO 19 I = 2, NCOBS
IF(START(I) - BEGTIM) 17, 17, 16
16 BEGTIM = START(I)
17 IF(END(I) - ENDTIM) 18, 19, 19
18 ENDTIM = END(I)
19 CONTINUE
IF(BEGTIM - ENDTIM) 490, 490, 405
C
C NO OVERLAP PERIOD FOUND
C
405 MM = NCOBS + 1
INDIC = + 1
NOREAD = NOREAD + 1
READ(19,END=8998) STTIM, FINTIM, NUMST, DATYP, NDGREE, COEFF,
11SAT, ROBWT, SOFT, STNAME, FICRP, IOCI
C
C THIS MEANS NO GATING IS DONE BY CODS, NOTICE THIS IS GOOD FOR
C RANGE ONLY, DO NOT USE IF RANGE RATE IS ALSO USED
C
111(2) = 111(2) + 2
START(MM) = STTIM
END(MM) = FINTIM
TYPE(MM) = DATYP
NOMST(MM) = NUMST
STATNM(MM) = STNAME
NDEG(MM) = NDGREE
STDFIT(MM) = SOFT
WTOB(MM) = ROBWT
KK = NDEG(MM) + 1
DO 451 J = 1, KK
451 A(MM,J) = COEFF(J)
452 IF(J - 13) 453, 450, 450
453 J = J + 1
A(MM,J) = 0.
GO TO 452
450 WRITE(6,5001) MM, START(MM), END(MM), NOMST(MM), TYPE(MM),
INDEG(MM), (A(MM,J), J = 1, 13)
DO 435 I = 1, MM
IF(ENC(I) - ENDTIM) 431, 431, 435

```

```

431 START(I) = START(MM)
    END(I)   = END(MM)
    TYPE(I)  = TYPE(MM)
    NOMST(I) = NOMST(MM)
    STATNM(I) = STATNM(MM)
    NDEG(I)  = NDEG(MM)
    STDFIT(I) = STDFIT(MM)
    WTOB(I)  = WTOB(MM)
    DO 432 J = 1, 13
432 A(I,J) = A(MM,J)
    GO TO 415
435 CONTINUE
490 IF(NOREAD - (2 * NOCBS)) 251, 210, 210
251 DO 252 III = 1, NOCBS
252 NSTAT(III) = NOMST(III)
210 IF(INDIC) 253, 215, 215
215 DO 250 JJJ = 1, NOOBS
    DO 216 III = 1, NOOBS
        IF(NCMST(III).EQ.NSTAT(JJJ)) GO TO 220
216 CONTINUE
220 IF(III.EQ.JJJ) GO TO 25C
    TEMP20 = START(JJJ)
    TEMP21 = END(JJJ)
    TEMP22 = TYPE(JJJ)
    TEMP23 = NOMST(JJJ)
    TEMP24 = STATNM(JJJ)
    TEMP25 = NDEG(JJJ)
    TEMP26 = STDFIT(JJJ)
    TEMP27 = WTOB(JJJ)
    DO 270 KKK = 1, 13
270 STOR(KKK) = A(JJJ, KKK)

    START(JJJ) = START(III)
    END(JJJ)   = END(III)
    TYPE(JJJ)  = TYPE(III)
    NOMST(JJJ) = NOMST(III)
    STATNM(JJJ) = STATNM(III)
    NDEG(JJJ)  = NDEG(III)
    STDFIT(JJJ) = STDFIT(III)
    WTOB(JJJ)  = WTOB(III)
    DO 277 KKK = 1, 13
277 A(JJJ, KKK) = A(III, KKK)
    START(III) = TEMP20
    END(III)   = TEMP21
    TYPE(III)  = TEMP22
    NOMST(III) = TEMP23
    STATNM(III) = TEMP24
    NDEG(III)  = TEMP25
    STDFIT(III) = TEMP26
    WTOB(III)  = TEMP27
    DO 276 KKK = 1, 13
276 A(III, KKK) = STOR(KKK)
250 CONTINUE
    DO 260 I = 1, NK
260 WRITE(6,5001) I, START(I), END(I), NOMST(I), TYPE(I), NDEG(I),
    1(A(I,J), J = 1, 13)
    IF(INCCT) 139, 139, 253
253 I = BEGTIP
    M = 1
14 N = NOOBS
    IF(NOPRST) 25, 15, 15
15 IF(NCMST(N) - NOPRST) 2C, 25, 20
20 N = N - 1
    IF(N) 30, 30, 15
30 WRITE(6,5002)

```

```

5002 FORMAT(1H0, 50X, 33F NO TWO WAY RANGING STATION FOUND//)
      GC TO 9C00
25  NUMST = NOMST(N)
      NAMBR(N) = NAMB(NUMST)
      TEMP3 = NAMB(NUMST) * TEMP2
      TD = T + .5 * TEMP3
      NUMST = NOMST(N)
      STNCEL = STADEL(NUMST)
      KOUNT = 1
      N2WAY = N
      GO TO 80
31  CALL CHEBY(ND, AC, TD, Y, XST, XEND)
      Y = Y * 1.0D-C6
      Y2 = Y
      YSAVE = Y
      TEMP1 = .5 * Y / 8640C.
      DELCOD(M,N) = TEMP1
      TP = TD - .5 * TEMP3 + TEMP1
      IF(DABS(T / TP - 1.C) - EPSLON) 40, 40, 35
35  TD = TD - TEMP1
      GO TO 31
40  DELAY(N) = Y + NAMB(NUMST) * DELA - STNCEL
      DEL2W = Y + NAMB(NUMST) * DELA - STNCEL
      DELCUT(M,N) = Y - STNCEL
      D2WAY = Y - STNCEL
      WRITE(6,NAM1)
      T = TP
      IF(NOPRST) 42, 41, 41
42  IF(NOPTS) 26, 26, 65
26  N = N - 1
      IF(N) 65, 65, 25
41  N = NCOBS
43  IF(NOMST(N) - NOPRST) 50, 45, 50
45  N = N - 1
      IF(N) 65, 65, 43
50  TEMP3 = 0.5 * TEMP2 * (NAMB(NOPRST) + NAMB(NUMST))
      TD = T + TEMP3
      NUMST = NCMST(N)
      NAMBR(N) = NAMB(NUMST)
      STNCEL = STADEL(NUMST)
      KOUNT = 2
80  NSAV = N
      IF(NCPTS) 77, 77, 76
76  N = 1
77  ND = NDEG(N)
      XST = START(N)
      XEND = END(N)
      KK = ND + 1
      DO 81 K = 1, KK
81  AC(K) = A(N,K)
      N = NSAV
      GO TO (31, 51), KOUNT
51  CALL CHEBY(ND, AC, TD, Y, XST, XEND)
      Y = Y * 1.0D-C6
      Y4 = Y
      YSAVE = Y
      TEMP1 = .5 * Y / 8640C.
      DELCOD(M,N) = TEMP1
      IF(Y4 - Y2) 52, 52, 55
52  TD = TD + .5 * TEMP2
53  CALL CHEBY(ND, AC, TD, Y, XST, XEND)
      Y = Y * 1.0D-C6
      YSAVE = Y
      Y4 = Y
      DELAY(N) = Y + (NAMB(NUMST) + NAMB(NOPRST)) * DELA - STNCEL + DELA
      TEMP1 = .5 * Y / 8640C.
      DELCOD(M,N) = TEMP1
      WRITE(6,NAM3)
      IF(Y4 - Y2) 54, 54, 55

```

```

54 TP      = TD - TEMP3 + TEMP1 - .5 * TEMP2
GO TO 56
55 TP      = TD - TEMP3 + TEMP1
  DELAY(N) = Y + (NAMB(NUMST) + NAMB(NCPRST)) * DELA - STNDEL
56 IF(CABS(T / TP - 1.C) - EPSLON) 60, 60, 57
57 TD      = TC - TEMP1
GO TO 51
60 DELAY(N) = DELAY(N) - DEL2W
  DELCUT(M,N) = Y - D2WAY - STNDEL
  N          = N - 1
  IF(N) 65, 65, 43
65 IF(NCPTS) 82, 82, 83
82 DC 66 N = 1, NOCBS
  R(N)     = DELAY(N) * C / (2.0 * REARTH)
  RTVI(M,N) = R(N) * REARTH * 100C.0
  TVTIME(M,N) = T
66 ROBS(N) = R(N) * R(N)
GO TO 85
83 R(JJ)   = DELAY(N) * C / (2.0 * REARTH)
  RTVI(M,N) = R(JJ) * REARTH * 100C.0
  TVTIME(M,N) = T
85 IF(NCPTS) 90, 90, 145
90 IF(NCOBS - 3) 145, 91, 91
91 I       = 1
  NLSCIT   = 1
  WT(1)    = - 1.0
  CALL READCL(SYRE, SCNTH, SDAYY, SHOURS, SUETS, SSECCN, KK, T, SE)
  KYR      = SYRE
  KMON     = SONTH
  KDAY     = SDAYY
  KHR      = SHOURS
  KMIN     = SUETS
  SECS     = SSECCN
  Y        = YP
202 IF(NCMST(I) - 47) 68, 67, 68
67 NOSTN   = 2
GO TO 75
68 IF(NCMST(I) - 58) 70, 69, 70
69 NOSTN   = 3
GO TO 75
70 IF(NCMST(I) - 66) 72, 71, 72
71 NOSTN   = 4
GO TO 75
72 IF(NCMST(I) - 68) 74, 73, 74
73 NOSTN   = 1
GO TO 75
74 WRITE(6, 5003) NOMST(I)
5003 FORMAT(1H0, 11HSTATION NO., 2X, 13, 2X, 58HIS NOT INCLUDED IN THE
1STATION CONSTANTS, PROGRAM HALTED.//18X, 65HCHECK THE STATION NO.,
20R ACC NEW STATION TO PROGRAM AND TRY AGAIN)
GC TC 9C00.
75 MDAY    = KDAY
  IF(KMON - IMON) 100, 110, 100
100 MDAY    = MDAY + NDAY(IMON)
110 LROW    = MDAY - IDAY + 1
  ALAMCO    = ALAMDM(LRCW)
  ASECS     = 3600 * KHR + 60 * KMIN
  ASECS     = ASECS + SECS
  DELT      = ASECS / 3600.C+00
  ALAMCE    = STACON(NOSTN, 1)
  THETAD    = STACON(NOSTN, 2)
  ALT       = STACON(NOSTN, 3)
  SNTHED    = DSIN(THETAD)
  CSTHED    = DCCS(THETAD)
  CAPC      = CSTHED ** 2 + ((1. - F) * SNTHED) ** 2
  CAPC      = 1.C / DSQRT(CAPC)

```

```

CAPS = CAPC * (1.0 - F) ** 2
ALT = ALT * 4.77865E-08
TEMP = (CAPS + ALT) / (CAPC + ALT)
TEMP = TEMP * SNTHED / CSTHED
THETAG = DATAN(TEMP)
SNTHEG = DSIN(THETAG)
CSTHEG = DCOS(THETAG)
DELTA = ALAMDO + ALAMDE + DELT * 2.625161333D-01
SDELTA = DSIN(DELTA)
CDELTA = DCOS(DELTA)
RHOCAP = (CAPS + ALT) ** 2
RHOCAP = RHOCAP * SNTHED ** 2 + ((CAPC + ALT) * CSTHED) ** 2
RHOCAP = DSQRT(RHOCAP)

C
C
C
COMPUTATION OF INERTIAL CCORDINATES OF OBSERVATION POINT

XM = RHOCAP * CSTHEG * CDELTA
YM = RHOCAP * CSTHEG * SDELTA
ZM = RHOCAP * SNTHEG
XT(1) = XM
YT(1) = YM
ZT(1) = ZM
I = I + 1
IF(I - NOOBS) 202, 202, 116
116 I = 1
IF(NOITER - 1) 118, 118, 400
118 CALL COMP(XT, YT, ZT, RCBS, XOF, YOF, ZOF)
X = XCF
Y = YOF
Z = ZOF
WRITE(6,500C) F, TOL, DELA, NOOBS, NOPRST, DELMIN, REARTH, C
5000 FORMAT(1H1, 50X, 25H ATX EXPERIMENTAL RANGING//54X,
116H INPUT PARAMETERS//4X, 20HEARTH'S FLATTENING = E15.8, 5X,
223H CONVERGENCE TOLERANCE = E15.8, 5X, 6H DELA = E15.8,
32X, 3H SEC//4X, 27H NO. OF OBSERVING STATIONS = 15, 5X,
422H NC. OF PRIME STATION = 15, 5X, 8H DELTAT = E15.8, 2X, 4H MINS//
518H RADIUS OF EARTH = E20.8, 2X, 3H KM., 5X, 16H SPEED OF LIGHT =
6E20.8, 2X, 10H KM. / SEC.//)
WRITE(6,5010) X, Y, Z
5010 FORMAT(1H0,
166X, 3HX = E15.8//1CX,
259H INITIAL ESTIMATE OF INERTIAL COORDINATES OF SATELLITE Y =
3 E15.8//66X, 3HZ = E15.8)
WRITE(6, 5005)
5005 FORMAT(1H1, 50X, 6H OUTPLT//
1 9X, 4H TIME, 26X, 19H SUB-SATELLITE POINT//1X,
222H YR MO DAY HR MIN SEC, 6X, 13H LONGITUDE(+E), 2X,
324H LATITUDE(+N) HEIGHT(KM)//)
400 CLMQI(1,1) = X - XT(1)
CLMQI(1,2) = Y - YT(1)
CLMQI(1,3) = Z - ZT(1)
RCOMP = CLMQI(1,1) ** 2 + CLMQI(1,2) ** 2 + CLMQI(1,3) ** 2
RC2 = DSQRT(RCOMP)
DELR(1, 1) = .5 * (ROBS(1) - RCOMP)
IF(I - NOOBS) 120, 130, 130
120 I = I + 1
GO TO 400
130 CALL GLSP(CLMQI, NOCBS, 3, DELR, 1, DLTAQI, RESID, SUM, WT, 0,
1STDERR)
X = X + DLTAQI(1,1)
Y = Y + DLTAQI(2,1)
Z = Z + DLTAQI(3,1)
XS = X * X
YS = Y * Y
ZS = Z * Z
NLSQIT = NLSQIT + 1
SUM = 0.
DO 300 I = 1, NOCBS

```

```

300 SUM      = SUM + (.2.C * CELR(I,1)) ** 2
    FIT      = DSQRT(SUM / NCOBS)
    I = 1
    IF(NLSQIT - 20) 350, 35C, 8C00
350 IF(FIT - TOL) 500, 500, 400
C
C      BEGIN COMPUTATION OF SUB-SATELLITE POINT
C
500 RHOS      = XS + YS + ZS
    I          = 0.0
    DTHEG      = 0.000
    TES        = F * (2.000 - F)
    T2         = (1. - F) ** 2
    RHO        = DSQRT(RHOS)
    SNTHEG     = Z / RHO
    CSTHEG     = DSQRT((XS + YS) / RHOS)
    TEMP1      = RHO * CSTHEG
    CSDELTA    = X / TEMP1
    SNDELTA    = Y / TEMP1
    DELTA      = DATAN2(SNDELTA, CSDELTA)
    XLONG      = DELTA - ALAMDO -.2625161333D+00 * DELT
    DEL        = DATAN(Z/DSQRT(XS + YS))
501 THETAG    = DEL - DTHEG
    I          = I + 1
    T10        = DCOS(THETAG)
    T11        = T10 * T10
    RC         = DSQRT((1.000 - TES) / (1.000 - TES * T11))
    XLAT       = DATAN(DTAN(THETAG) / T2)
    T12        = XLAT - THETAG
    T13        = DSIN(T12)
    T14        = T13 * T13
    T15        = DCOS(T12)
    HT         = DSQRT(RHOS - RC * RC * T14) - RC * T15
    T20        = HT * T13 / RHO
    DTHEG      = DARSIN(T20)
    IF(I.LT.10) GO TO 5C1
    XLAT       = XLAT * RADDEG
    XLONG      = XLONG * RADDEG
    HT         = HT * AE
520 IF(DABS(XLONG) - 18C.0) 510, 510, 700
700 IF(XLONG) 701, 702, 702
701 XLONG      = XLONG + 360.0
    GO TO 510
702 XLONG      = XLONG - 360.0
510 WRITE(6,5004) KYR, KMCN, KDAY, KHR, KMIN, SECS, XLONG, XLAT, HT
5004 FORMAT(1H0, I2, 1X, I2, 2X, I2, 1X, I2, 2X, I2, 2X, F5.2, 5X, F8.3,
    17X, F7.3, 4X, F10.2)
    WRITE(6,9035) X, Y, Z
9035 FORMAT(1H0, 4H X = E20.8//4H Y = E20.8//4H Z = E20.8)
145 T          = T + DELTAT
    NOITER     = NOITER + 1
    YP        = Y
    M          = M + 1
    JJ        = JJ + 1
    IF(T - ENDTIM) 150, 150, 9000
150 IF(NOPTS) 14, 14, 25
9000 IF(NOPTS) 170, 170, 160
160 N          = N + 1
    JJ        = 1
    IF(N - NCOBS) 5, 5, 165
165 N          = 1
170 NTOT       = M - 1
    WRITE(9,1000) NTOT
    WRITE (6,1000) NTOT
1000 FORMAT(I5)
    M          = 0

```

```

8999 M      = M + 1
      WRITE(9,1001)(RTV1(M,NN), NN = 1, NOOBS)
      WRITE(6,1001)(RTV1(M,NN), NN = 1, NOOBS)
      WRITE(9,1001)(TVTIME(M,NN), NN = 1, NOOBS)
      WRITE(6,1001)(TVTIME(M,NN), NN = 1, NOOBS)
1001 FORMAT(3D25.15)
      IF(M - NTOT) 8999, 9010, 9010
9010 DO 9050 M = 1, NTCT
      DO 9050 L = 1, NOCBS
C
C      CONVERT RANGE MEASUREMENTS TO DUL UNITS
C
      IOBN      = IOBN + 1
      ROUT      = RTV1(M,L) * 1.0C-07
      TIME      = TVTIME(M,L)
      T30       = (NAMBPRL) * DELA) / (2.000 * 86400.000)
      TIME      = TIME + T30 + CELDOD(M,L)
      IF(L.NE.N2WAY) TIME = TIME + T30 - RTV1(M,N2WAY) / (1000.0 * C *
1      86400.0)
      CALL READCL(SYRE, SCNTH, SDAYY, SHOURS, SUETS, SSECON, KK, TIME,
1SE)
      KYR        = SYRE
      KMON       = SCNTH
      KDAY       = SDAYY
      KHR        = SHOURS
      KMIN       = SUETS
      SECS       = SSECON
      IF(CFLOAT(KYR/4) - DFL0AT(KYR) / 4.000) 94, 92, 94
92 NCAY(2) = 29
94 NODAY      = 0
      LL        = KMON - 1
      DO 95 II = 1, LL
95 NCDAY      = NCCAY + NCAY(II)
      NODAY     = NODAY + KDAY
      NSJCS     = (KYR - 58) * 365 + ((KYR - 1) / 4 - 14) + NCDAY + 104
      SJDS      = 100.00 * (DFLCAT(NSJDS) + DFL0AT(KHR) / 24.00
1      + DFL0AT(KMIN) / 1440.00 + SECS / 86400.00)
      OBCV=0.00
      RCBWT     = WTCB(L)
      STNAME    = STATNM(L)
      SDFT      = STDFIT(L)
9050 WRITE (29)SJDS,STNAME,RCUT ,TRANCM,FIDRP,OBCV,ISAT,IOBN,ROBWT,COC,
1POC,SDFT,FNRRES,TMER,TCRV,IESN,ICRPN,IOT,IEF,ITTI,ITCI,IOCI,IOCI
      GO TO 5
8000 WRITE(6,3050)
3050 FORMAT(1H0, 45X, 31F PROGRAM HALTED, NO CONVERGENCE)
8598 REWIND 29
      WRITE (6,1300)
1300 FORMAT ('1OUT',19X,'STNAME   OBSV',16X,'FEEDBACK   CORR VAL',
17X,'SAT NO,OB NO, OB WT',8X,'ST. DEV.   CBTP ECFG OCI')
      DO 1302 J=1,32000
      READ (29,END=5210)SJDS,STNAME,ROUT,TRANOM,FIDRP,OBCV,ISAT,IOBN,
1OBWT,COC,POC,SDFT,FNRRES,TMER,TCRV,IESN,ICRPN,IOT,IEF,ITTI,ITCI,
210CI
1302 WRITE (6,1301)SJDS,STNAME,ROUT,I11(1),I11(2),OBCV,ISAT,IOBN,
1OBWT,SDFT,IOT,IEF,ICCI
1301 FORMAT (1X,D20.12,1X,A8,1X,D20.12,1X,2I5,1X,D14.6,1X,17, 16,
11X,C12.4,1X,C12.4,1X,I2,1X,I2,1X,I6)
5210 STOP
      END
/*

```

0583 CARDS

C
C
C

SUBROUTINE COMP

```

SUBROUTINE COMP(XT, YT, ZT, ROBS, XOF, YOF, ZOF)
  IMPLICIT REAL*8(A-H,O-Z,$)
  REAL*8 LAMBDA
  DIMENSION X(3),Y(3),Z(3),XO(2),YO(2),ZO(2),RO(3,2),RO2(2),CHECK(2)
  1, XT(10), YT(10), ZT(10), ROBS(10), R(3,3), RHO(3), R2(3),
  2 RHON(3,3)
  DO 25 I=1,3
    R(1,I) = -XT(I)
    R(2,I) = -YT(I)
    R(3,I) = -ZT(I)
    RHO(I) = DSQRT(ROBS(I))
    X(I)=R(1,I)
    Y(I)=R(2,I)
  25 Z(I)=R(3,I)
  DO 35 I=1,3
  35 R2(I)=DOT(R(1,I),R(1,I))
    E21=.5D0*(RHO(2)**2-RHO(1)**2-(R2(2)-R2(1)))
    E31=.5D0*(RHO(3)**2-RHO(1)**2-(R2(3)-R2(1)))
    DELTA1=(Z(3)-Z(1))*(Y(2)-Y(1))-(Z(2)-Z(1))*(Y(3)-Y(1))
    A=((X(2)-X(1))*(Y(3)-Y(1))-(X(3)-X(1))*(Y(2)-Y(1)))/DELTA1
    B=(E31*(Y(2)-Y(1))-E21*(Y(3)-Y(1)))/DELTA1
    DELTA2=-DELTA1
    C=((X(2)-X(1))*(Z(3)-Z(1))-(X(3)-X(1))*(Z(2)-Z(1)))/DELTA2
    D=(E31*(Z(2)-Z(1))-E21*(Z(3)-Z(1)))/DELTA2
    EPS1=A*A+C*C+1.D0
    EPS2 =2.D0*(A*B+C*D+X(1)+C*Y(1)+A*Z(1))
    EPS3 =B*B+D*D+2.D0*D*Y(1)+2.D0*B*Z(1)+R2(1)-RHO(1)**2
    RAD =EPS2 **2-4.D0*EPS1*EPS3
    XO(1)=(-EPS2+DSQRT(RAD))/(2.D0*EPS1)
    XO(2)=(-EPS2-DSQRT(RAD))/(2.D0*EPS1)
    DO 30 I=1,2
    YO(I)=C*XO(I)+D
    ZO(I)=A*XO(I)+B
    RO(1,I)=XO(I)
    RO(2,I)=YO(I)
    RO(3,I)=ZO(I)
    XOF=XO(1)
    YOF=YO(1)
    ZOF=ZO(1)
    DO 75 K=1,3
    DO 75 J=1,3
  75 RHON(J,K)=R(J,K)+RO(J,I)
    IF(DOT(RHON,R)/(DSQRT(R2)*RHO(1)).LE.0.D0) RETURN
  30 CONTINUE
  RETURN
  END

```

CDOTT

```

FUNCTION DOT(A,B)
  IMPLICIT REAL*8(A-H,O-Z,$)
  DIMENSION A(3),B(3)
  DOT=0.D0
  DO 1 I=1,3
  1 DOT=DOT+A(I)*B(I)
  RETURN
  END

```

0058 CARDS


```

C
C
C
SUBROUTINE CHEBY
SUBROUTINE CHEBY(ND,C,XP,Y,XF,XL)
IMPLICIT REAL*8(A-4,J-2)
DIMENSION C(13)
X= 2.00*(XP-XF)/(XL-XF) -1.00
TO=1.00+0
T1=X
Y=C(1)*TO+C(2)*T1
IF (ND.LT.2) RETURN
NT=ND+1
DO 10 K=3,NT
TN=2.0*X*T1-TO
Y=Y+C(K)*TN
TO=T1
10 T1=TN
RETURN
END

```

0019 CARDS

```

C
C
C
SUBROUTINE DOT
IMPLICIT REAL*8(A-4,J-2,S)
DIMENSION A(3),B(3)
DOT=0.00
DO 1 I=1,3
1 DOT=DOT+A(I)*B(I)
RETURN
END
//OBJECT DD *
/*
//OBJECT DD *

```

0013 CARDS

```

C
C
C
SUBROUTINE READCL
SUBROUTINE READCL (SYRE,SONTH,SDAYY,SHOURS,SUETS,SSECON,
1KK,SSUNA,SE)
IMPLICIT REAL*8(A-4,J-2,S)
DIMENSION YM(28),YJ(28),OM(13)
READ CALENDAR DAYS
DATA OM/31.00,28.00,31.00,30.00,31.00,30.00,31.00,31.00,30.00,
131.00,30.00,31.00,0.00/
DATA YM/58.00,59.00,60.00,61.00,62.00,63.00,64.00,65.00,66.00,
167.00,68.00,69.00,70.00,71.00,72.00,73.00,74.00,75.00,76.00,
277.00,78.00,79.00,80.00,81.00,82.00,83.00,84.00,85.00/
C
SAD MEAN JULIAN DATE
DATA YJ/36203.00, 36568.00,36933.00,37299.00,37664.00,38029.00,
138394.00,38760.00,39125.00,39490.00,39855.00,40221.00,40586.00,
240951.00,41316.00,41682.00,42047.00,42412.00,42777.00,43143.00,
343508.00,43873.00,44238.00,44604.00,44969.00,45334.00,45699.00,
446065.00/
SUNA=SSUNA
E=+1.00
I=1
K=1
66 IF (YJ(1)-SUNA)167,67,12
12 WRITE (6,301) SUNA ,YJ(1),YJ(28)
STOP
301 FORMAT (' OUTSIDE CALENDAR RANGE, MJD =',F20.4,' RANGE =',
1F20.5,' TO ',F20.5)

```

FCP70010

FCP70020
FCP70030
FCP70040
FCP70070

67	IF (YJ(28)-SUNA)12,68,68	FCP70080
68	I=2	FCP70090
71	IF (YJ(1)-SUNA+.500)69,70,70	FCP70100
69	I=I+1	FCP70110
	GO TO 71	
70	YRE=YM(I-1)	FCP70130
100	SUNM=SUNA-YJ(I-1)	FCP70140
	KSUNM=SUNM	FCP70150
	SUNM1=KSUNM	FCP70160
	IYRE=YRE+.0100	FCP70170
	IF(MOD(IYRE,4)) 77,78,77	
78	OM(2)=29.00	FCP70190
	GO TO 105	
77	OM(2)=28.00	FCP70210
105	I=1	FCP70220
	SM=OM(I)+1.00	FCP70230
73	IF (SM-SUNM)81,81,72	FCP70240
81	I=I+1	FCP70250
	IF (I-13)210,72,210	FCP70260
210	SM=SM+OM(I)	FCP70270
	GO TO 73	
72	SD=SUNM-(SM-OM(I))+1.00	FCP70290
	DNTH=I	FCP70300
201	KSD=SD	FCP70310
	DAC = SD	
	DAYY=KSD	FCP70320
	SDAC=DAC	
	SDAC=IDINT(SDAC)	
	DDAC=SDAC	
	SD=DAC-DDAC	
	IF (SD)220,221,221	FCP70340
220	SD=1.00+SD	FCP70350
221	HOURL=SD*24.00	FCP70360
	KHOURL=HOURL	FCP70370
	HOURS=KHOURL	FCP70380
	UE=(HOURL-HOURS) *60.00	FCP70390
	KUE=UE	FCP70400
	UETS=KUE	FCP70410
	SECON=(UE-UETS)*60.00	FCP70420
	IF (DNTH)207,208,207	FCP70430
208	DNTH=12.00	FCP70440
207	CONTINUE	
99	SYRE=YRE	
	SONTH=DNTH	
	SDAYY=DAYY	
	SHOURS=HOURS	
	SUETS=UETS	
	SSECON=SECON	
	SSUNA=SUNA	
	SE=E	
	RETURN	FCP71430
	END	0080 CARDS

C
C
C

SUBROUTINE GLSP

SUBROUTINE GLSP(A,MM,NV,B,IPP,X,U,SUM,WT,INVRS,STDERR)	005
IMPLICIT REAL*8 (A-H,O-Z)	
DIMENSION A (10,3), B (10,1), X (3,4), U (10,1), SUM (4),	
1 W (3,4), WT (10), STDERR (3)	
M=MM	008
N=NV	009
IP=IPP	010
INVRS=INVRS	011
AX=B,NORMAL EQUATIONS,A IS RECTANGULAR MATRIX,WITH M ROWS,N COLUMNS	012
WX=W1,NEW SET OF NORMAL EQUATIONS,W=TR(A)*A,W1= TR(A)*B..	013
TO FORM MATRIX W,WITH V ROWS AND V COLUMNS. (TR= TRANSPOSE)	014

C
C
C

C	WT = -1. MEANS NO WEIHTGS IN PRG. ASSIGN WEIGHTS = 1.	015
C	STDERR = STANDARD ERRORS OF UNKNOWNNS (X) .	016
C	INVRSE = 1. MEANS STD. ERRORS ARE CALCULATED.	017
C	INVRSE = 0. MEANS NO CALCULATIONS OF STD. ERRORS OF (X)*****	018
	KK=N+1	019
	IF(M-N) 80,83,81	020
81	DO 5 J=1,N	021
	DO 5 JJ=1,N	022
	W(J,JJ) = 0.	023
	IF (WT(1)) 611, 622, 622	
C	NO WEIGHTS CASE	025
611	DO 51 I= 1,M	026
51	W(J,JJ) = W(J,JJ) + A(I,J) * A(I,JJ)	027
	GO TO 5	028
C	WEIGHTS ARE PRESENT	029
622	DO 52 I=1,M	030
52	W(J,JJ) = W(J,JJ) + A(I,J) * A(I,JJ) * WT(I)	031
5	CONTINUE	032
C	TO FORM MATRIX W1 , WHICH IS STORED IN N+1 TO N+P COLUMNS OF W.	033
C	W1 ,WITH V ROWS AND IP COLUMNS.	034
	IF (INVRSE) 201,203,201	0035
203	KIP = IP	036
	GO TO 200	037
201	KIP = 1	038
200	DO 8 K = 1, KIP	039
	DO 7 J=1,V	040
	W(J,KK) = 0.	041
	IF (WT(1)) 612, 623, 623	
612	DO 511 I=1,M	043
511	W(J,KK) = W(J,KK)+A(I,J)* B(I,K)	044
	GO TO 7	045
623	DO 522 I= 1,M	046
522	W(J,KK) = W(J,KK)+A(I,J)* B(I,K) * WT(I)	047
7	CONTINUE	048
8	KK = KK + 1	049
	GO TO 106	050
83	DO 12 I=1,N	051
	IF (WT(1)) 613, 624, 624	
613	DO 512 J=1,N	053
512	W(I,J) = A(I,J)	054
	GO TO 12	055
624	DO 523 J=1,N	056
523	W(I,J) = A(I,J) * WT(I)	057
12	CONTINUE	058
	KKK = V+IP	059
	J=1	060
	IF (INVRSE) 401,403,401	061
403	KKK = KKK	062
	GO TO 400	063
401	KKK = KK	064
400	DO 13 JJ= KK,KKK	065
	IF (WT(1)) 614, 625, 625	
614	DO 513 I= 1,N	0067
513	W(I,JJ) = B(I,J)	068
	GO TO 13	069
625	DO 524 I= 1,N	070
524	W(I,JJ) = B(I,J) * WT(I)	071
13	J=J+1	072
106	IF (INVRSE) 71,72,71	073
71	K2 = N + 2	074
	KIP = V + IP	075
	KK = N + 1	076
	DO 75 I=1,N	077
	DO 75 J = K2,KIP	078
	IF (J - I - KK) 73,74,73	079

74	W(I,J) = 1.	080
	GO TO 75	081
73	W(I,J) = 0.	082
75	CONTINUE	083
72	CALL TRIANG (W,N,IP, DET)	084
C	WHERE W IS THE MATRIX WITH N ROWS AND(N+IP)COLUMNS TO BE TRIANGULA	085
C	RIZED	086
C	TO TEST THE SINGULARITY OF MATRIX W., W = TR(A)*A , W*X = W1.	087
	IF (DET) 82,80,82	088
82	CONTINUE	089
	GO TO 103	090
80	SUM(1) = -1.	092
	GO TO 104	093
103	CALL SOLVE (W,N,IP,X)	093
C	WHERE W IS THE TRIANGULARIZED, X IS THE SOLUTION MATRIX WITH	094
C	N ROWS AND IP COLUMNS.	095
C	TO COMPUTE RESIDUAL MATRIX , U = A*X - B.,U IS RESIDUAL MATRX(M,P)	096
105	DO 16 I=1,M	097
	DO 16 K=1,IP	098
	U(I,K) =0.	099
	DO 15 J=1,N	100
15	U(I,K) = U(I,K) + A(I,J) * X(J,K)	101
16	U(I,K) = B(I,K) - U(I,K)	102
	DO 18 K=1,IP	103
	SUM(K) = 0.	104
	DO 18 I=1,M	105
18	SUM(K) =SUM(K) +(U(I,K))**2	106
	IF (INVRS) 722,104,722	107
722	DO 723 I = 1,N	108
	DO 723 J= 2,IP	109
723	X(I,J) = X(I,J) * (SUM(1) / DFLOAT(M-N))	111
	DO 724 I = 1,N	112
	STDERR(I) = 0.	113
	II = I + 1	
724	STDERR(I) = DSQRT(X(I,II))	
104	RETURN	
	END	

0116 CARDS

C		
C	SUBROUTINE TRIANG	
C		
	SUBROUTINE TRIANG(A,M,N,DET)	120
	IMPLICIT REAL*8 (A-H,O-Z)	
	DIMENSION A (3,4)	
CTRIANG	TO BE USED WITH SUBROUTINE GLSP	119
C .	WHERE A IS THE MATRIX WITH M ROWS AND M+N COLUMNS TO	121
C	BE TRIANGULARIZED.	122
	DET=1.	124
	MM1=M-1	125
	MPN=M+N	126
	DO 5 J=1,MM1	127
	MAXX = J	128
	VALUE = DABS(A(J,J))	
	JPI=J+1	130
	DO 1 K=JPI,M	131
	IF (VALUE - DABS(A(K,J))) 2, 1, 1	
2	VALUE = DABS(A(K,J))	
	MAXX=K	134
1	CONTINUE	135
	DO 3 L=J,MPN	136
	TEMP = A(MAXX,L)	137
	A(MAXX,L) = A(J,L)	138
3	A(J,L) = TEMP	139
	IF (MAXX-J) 7,6,7	140
7	DET = - DET	141

6	DET = DET*A(J,J)	142
	ROW = A(J,J)	143
	DO 4 L=J,MPN	144
4	A(J,L) = A(J,L)/ROW	145
	DO 5 K=JPL,M	146
	ROW = A(K,J)	147
	DO 5 L=J,MPN	148
5	A(K,L) = A(K,L)-ROW*A(J,L)	149
	DET = DET*A(M,M)	150
	ROW = A(M,M)	151
	DO 9 L = 4,MPN	152
9	A(M,L) = A(M,L)/ROW	
	RETURN	154
	END	
		0041 CARDS
C		
C	SUBROUTINE SOLVE	
C		
	SUBROUTINE SOLVE(A,M,N,X)	159
	IMPLICIT REAL*8 (A-H,O-Z)	
	DIMENSION A (3,4), X (3,4)	
CSOLVE	TO BE USED WITH SUBROUTINE GLSP	158
C	WHERE A IS THE TRIANGULARIZED MATRIX WITH M ROWS AND	160
C	M+N COLUMNS.	161
C	X IS THE SOLUTION MATRIX WITH M ROWS AND N	162
C	COLUMNS.	163
	MM1=M-1	165
	DO 2 L=1,N	166
	MPL=M+L	167
	DO 2 K=1,MM1	168
	MMK=M-K	169
	MMKPL=MMK+1	170
	X(M,L)=A(M,MPL)/A(M,M)	171
	SUM=0.	172
	DO 3 I=MMKPL,M	173
3	SUM=SUM+A(MMK,I)*X(I,L)	174
2	X(MMK,L)=A(MMK,MPL)-SUM	175
	RETURN	176
	END	177
		0024 CARDS